## Periodic solutions to the relativistic Kepler problem: a dynamical systems approach

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### Talk Abstract

The motion of a relativistic particle in a Kepler potential can be described by the equation

$$\frac{\mathrm{d}}{\mathrm{d}t}\left(\frac{m\dot{x}}{\sqrt{1-|\dot{x}|^2/c^2}}\right) = -\alpha \, \frac{x}{|x|^3}, \quad x \in \mathbb{R}^2 \setminus \{0\},$$

where m > 0 is the mass of the particle, c is the speed of light, and  $\alpha > 0$  is a constant. Firstly, we illustrate the Hamiltonian formulation of the problem and we focus our attention on the description of the periodic and quasiperiodic solutions. Secondly, we deal with the perturbed equation

$$\frac{\mathrm{d}}{\mathrm{d}t}\left(\frac{m\dot{x}}{\sqrt{1-|\dot{x}|^2/c^2}}\right) = -\alpha \, \frac{x}{|x|^3} + \varepsilon \, \nabla_x U(t,x), \quad x \in \mathbb{R}^2 \setminus \{0\},$$

where U(t, x) is *T*-periodic in the first variable and  $\varepsilon \in \mathbb{R}$ . The analysis of the action-angle coordinates and an application of an higher dimensional version of the Poincaré–Birkhoff fixed point theorem allow to prove that, for  $\varepsilon$  small enough, the perturbed problem admits *T*-periodic solutions with prescribed winding number, bifurcating from invariant tori of the unperturbed problem. The talk is based on the paper [1] written in collaboration with Alberto Boscaggin and Walter Dambrosio.

**Keywords:** relativistic Kepler problem, periodic solutions, invariant tori, nearly integrable Hamiltonian systems, action-angle coordinates.

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# References

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